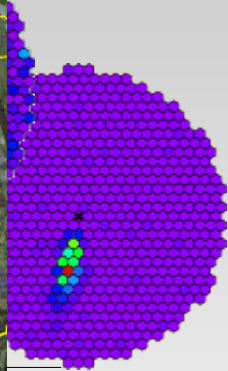
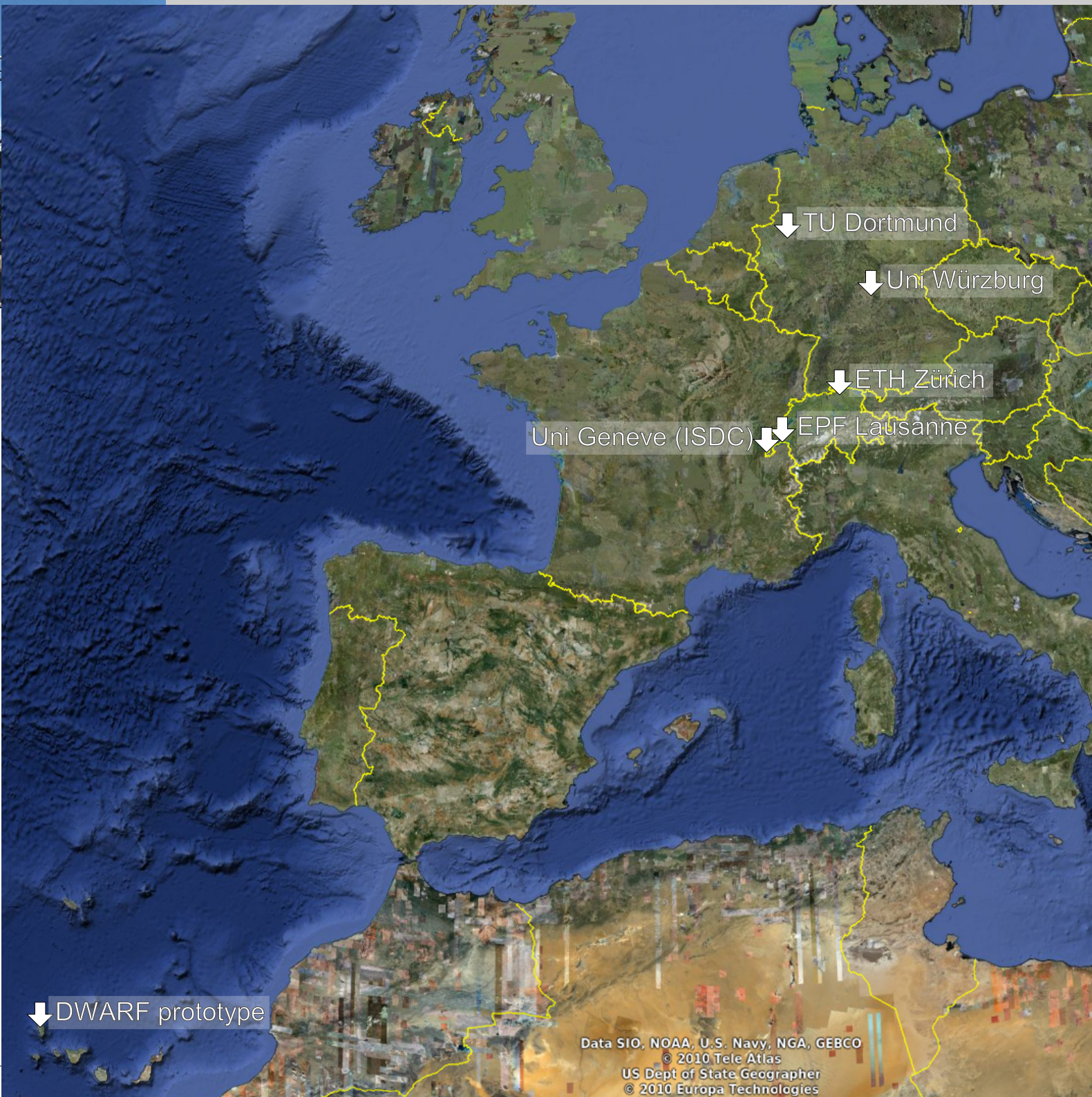


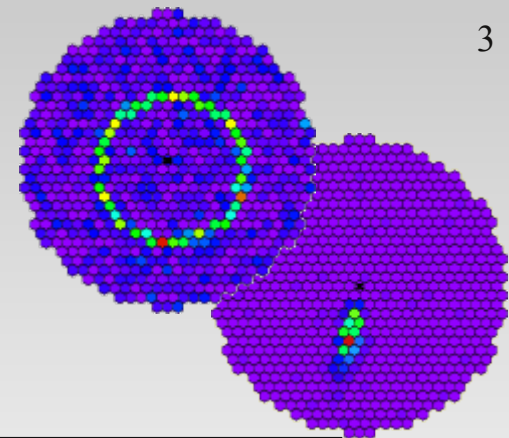
Status of FACT – The **F**irst **G**-**A**PD **C**herenkov **T**elescope

Thomas Bretz
for the FACT Collaboration





The Telescope



HEGRA CT3 at La Palma

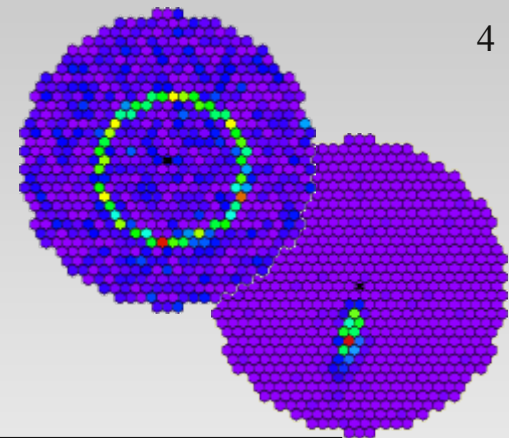


The idea

- Affordable monitoring telescope
 - observing only a few target objects (Mkn421, Mkn501, ...)
 - with large duty cycles
- Small-aperture as a prototype to a world-wide 24/7 monitoring network
- Refurbish one of the former HEGRA telescopes
 - switched off since ~2002
 - CT3 still at La Palma



The Telescope



Upgrade:

New drive system

- MAGIC type / installed

Refurbished mirrors

- hexagonal / larger area (9.5m^2)
- better reflectivity
- Extension with plastic mirrors (in prep.)

New camera:

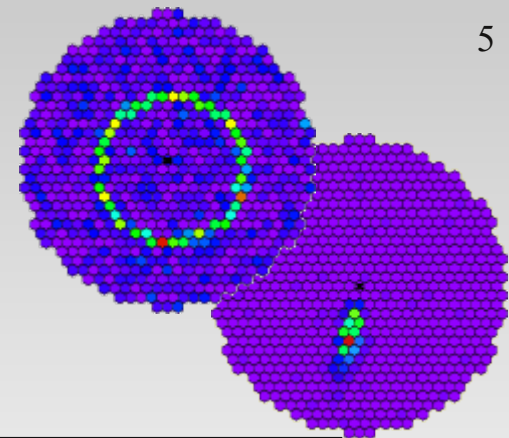
- First Cherenkov camera using G-APDs as photon detectors
- Integrated electronics

Artist view not long ago





The Telescope



Upgrade:

New drive system

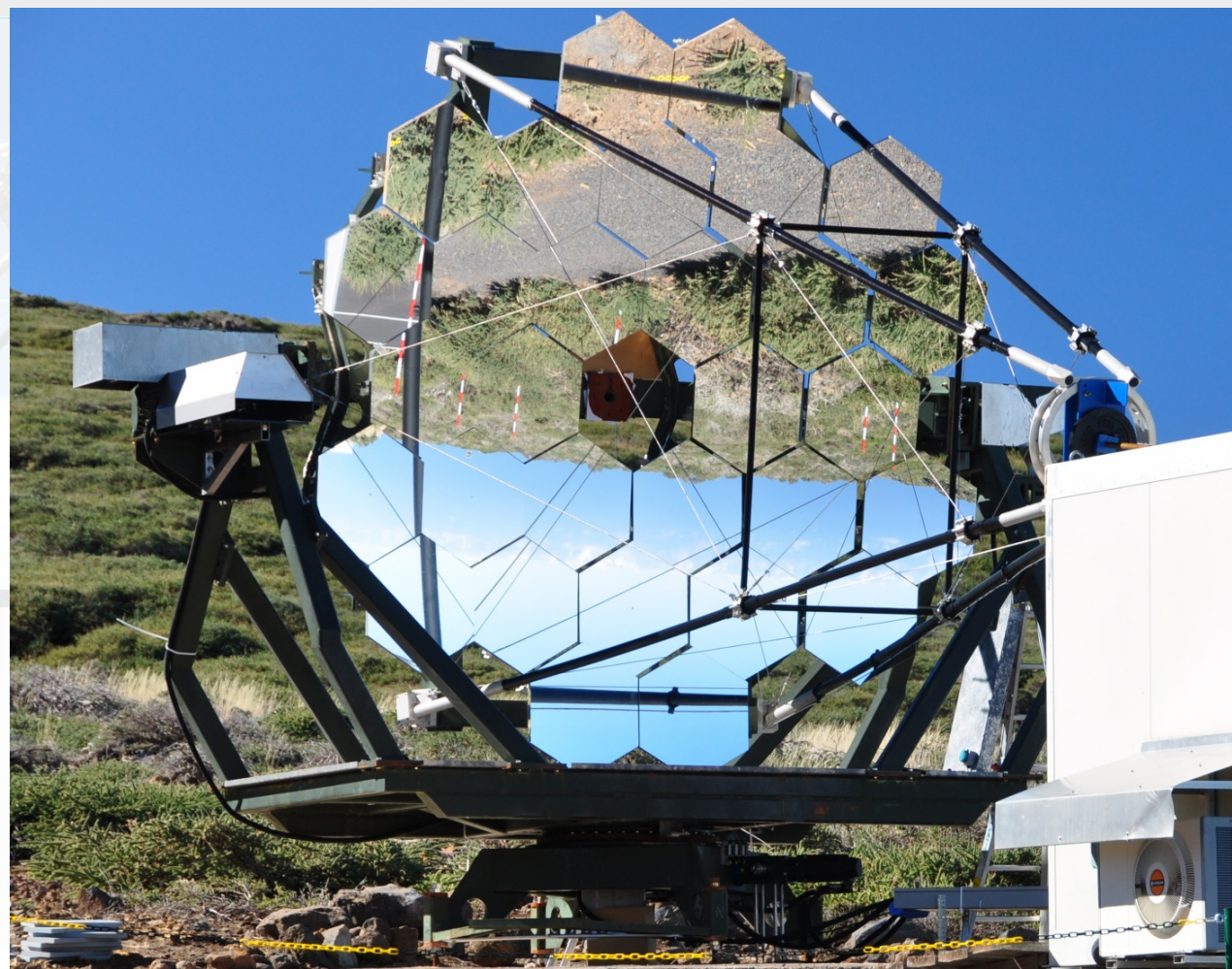
- MAGIC type / installed

Refurbished mirrors

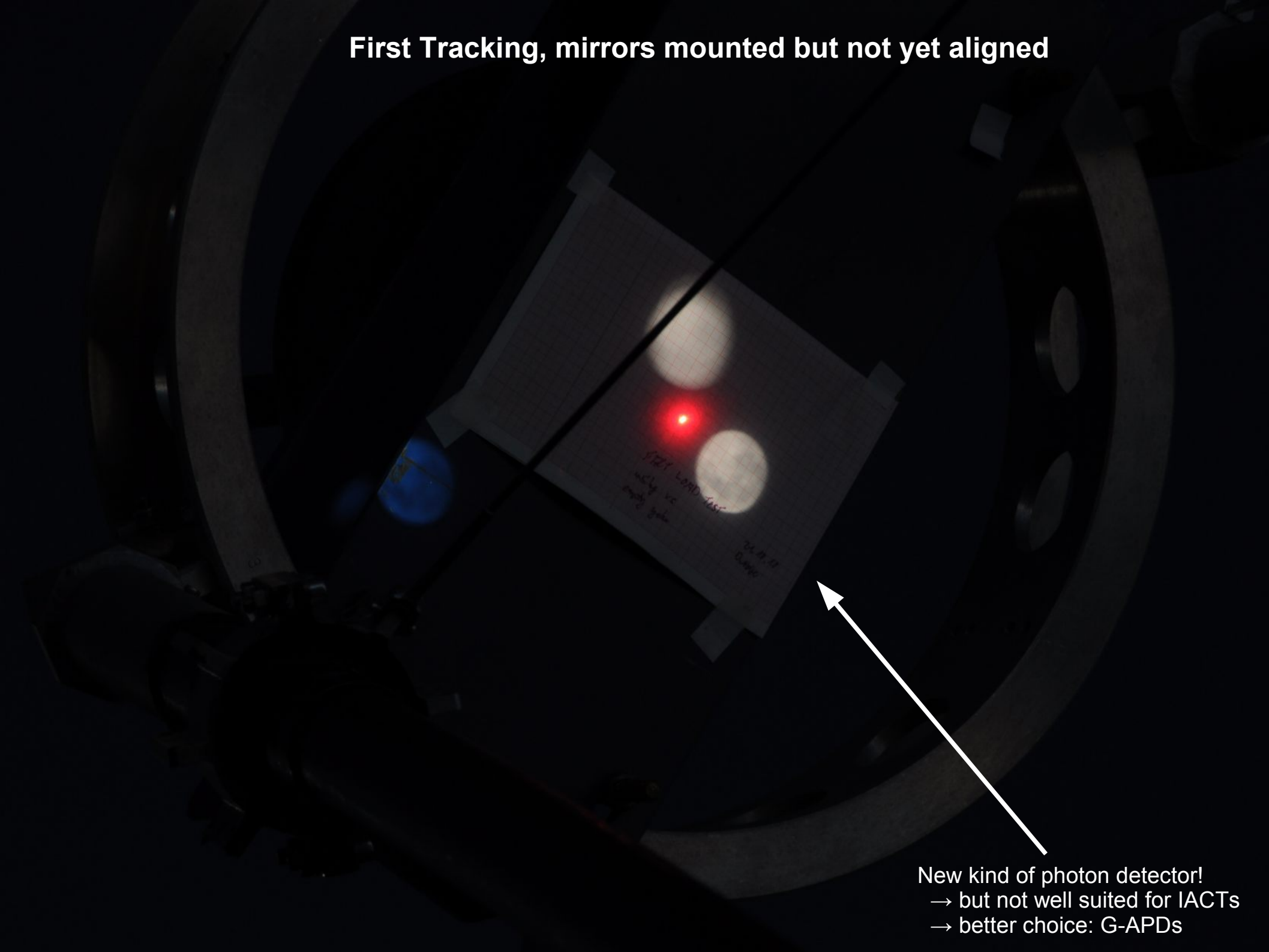
- hexagonal / larger area (9.5m^2)
- better reflectivity
- Extension with plastic mirrors (in prep.)

New camera:

- First Cherenkov camera using G-APDs as photon detectors
- Integrated electronics



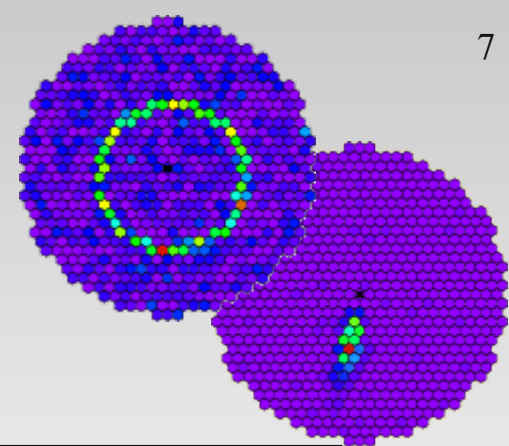
First Tracking, mirrors mounted but not yet aligned



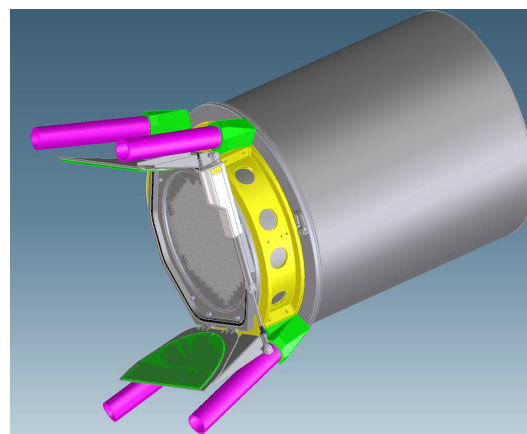
New kind of photon detector!
→ but not well suited for IACTs
→ better choice: G-APDs



The Camera

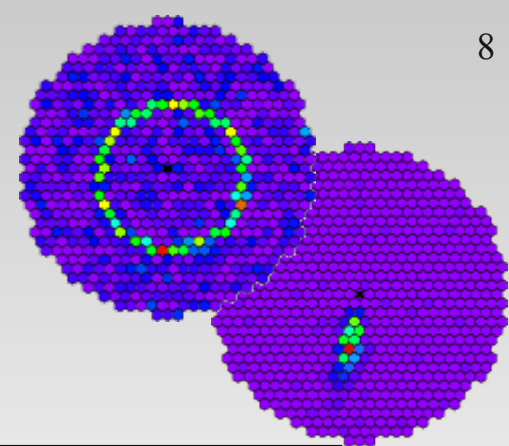


- First Cherenkov camera using G-APDs as photon detectors
- Integrated electronics
- 1440 pixels ($\sim 0.11^\circ/\text{pix}$)
- 4.5° field-of-view
- $\sim 100\text{kg}$
- DAQ: DRS4 (Ethernet readout)





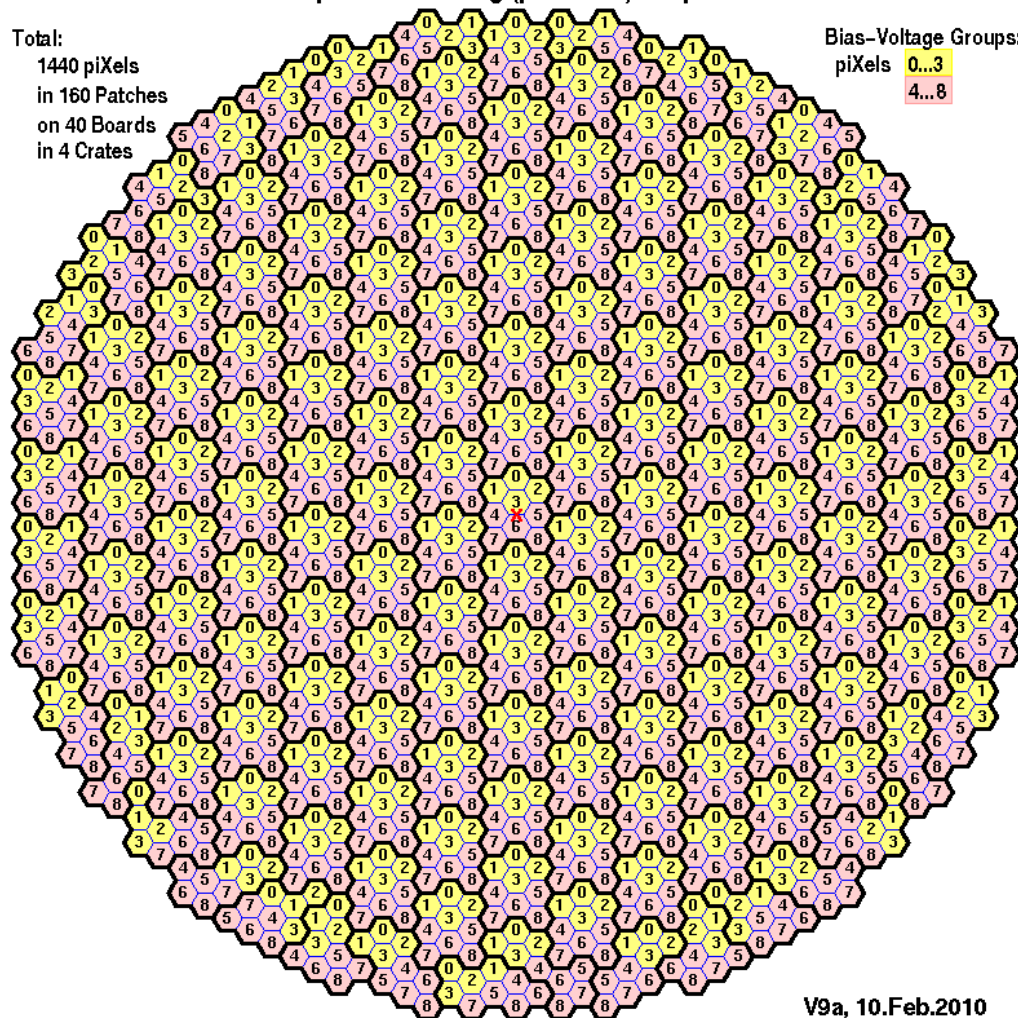
The Trigger



piXel Numbering (per Patch): cbpX

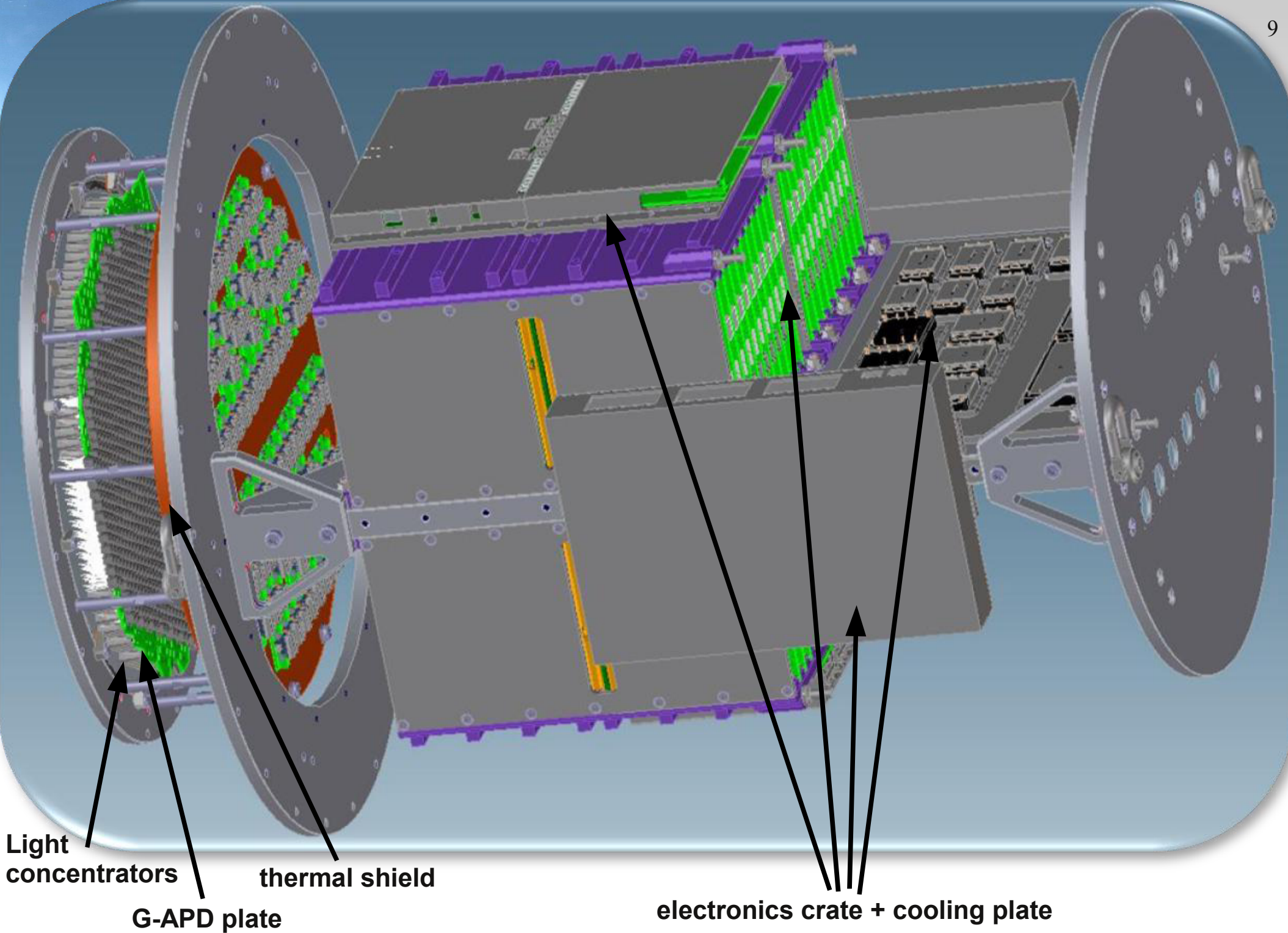
Total:
1440 piXels
in 160 Patches
on 40 Boards
in 4 Crates

Bias-Voltage Groups:
piXels 0...3
4...8



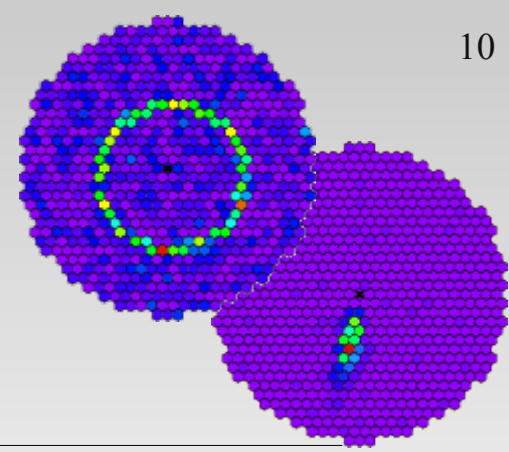
V9a, 10.Feb.2010

- First Cherenkov camera using G-APDs as photon detectors
- Integrated electronics
- 1440 pixels ($\sim 0.11^\circ/\text{pix}$)
- 4.5° field-of-view
- $\sim 100\text{kg}$
- DAQ: DRS4 (Ethernet readout)
- Sum-trigger with non-overlapping patches of nine pixels (close to *ideal*)
- Software trigger possible





Why G-APDs? advantages

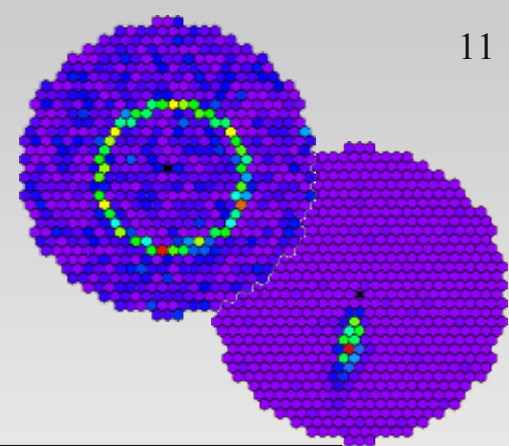


Most obvious advantages:

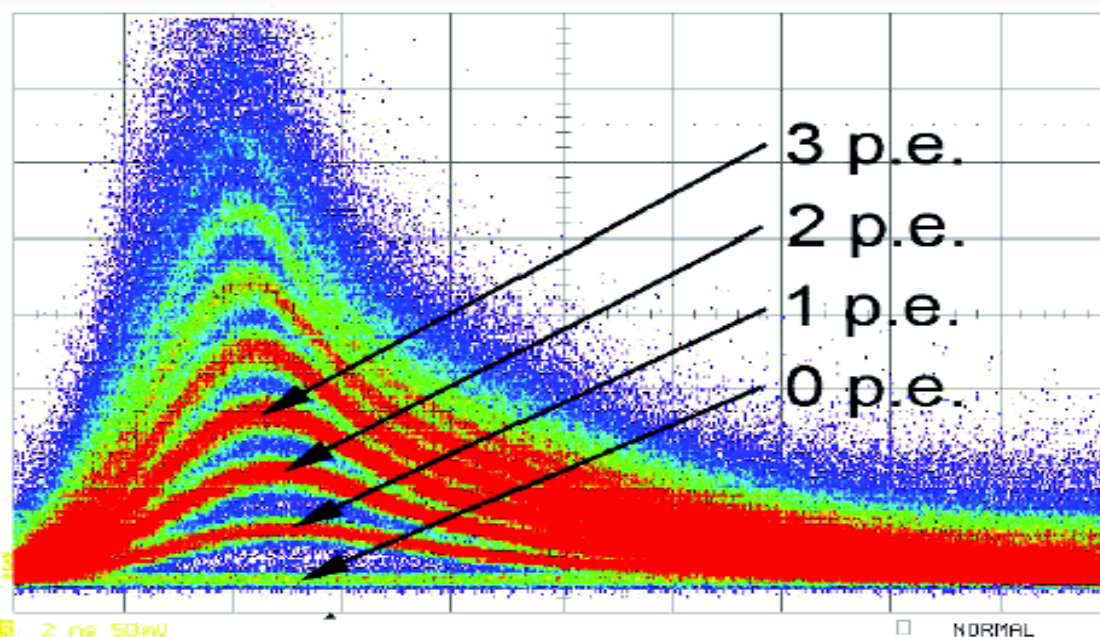
- Higher PDE (currently 30%-40%, but 70% possible)
- Single pe resolution (see next slide)
- Much lower bias voltage in the order of 70V
- More robust (can survive even exposure to direct sun-light)
- Easier to handle (less fragile)



Why G-APDs? single pe counting

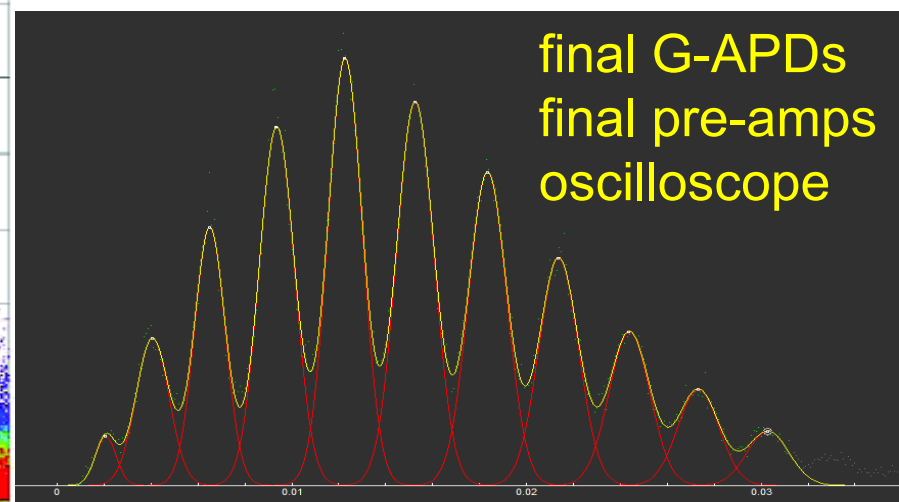


- G-APDs have very good single pe resolution
- Slower rise time than the best PMTs but very constant signals



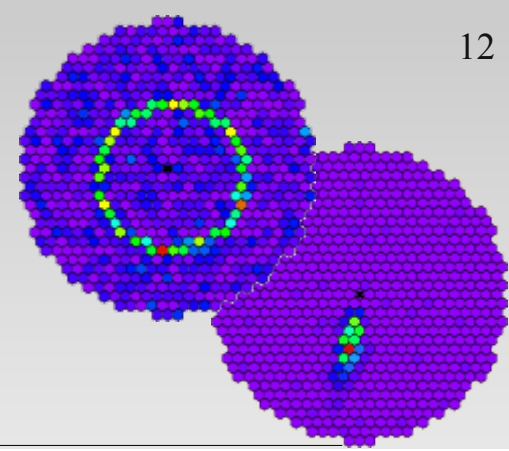
LED spectr., 2 ns horiz., 50 mV vert. spacing

Single pe separation using our electronics



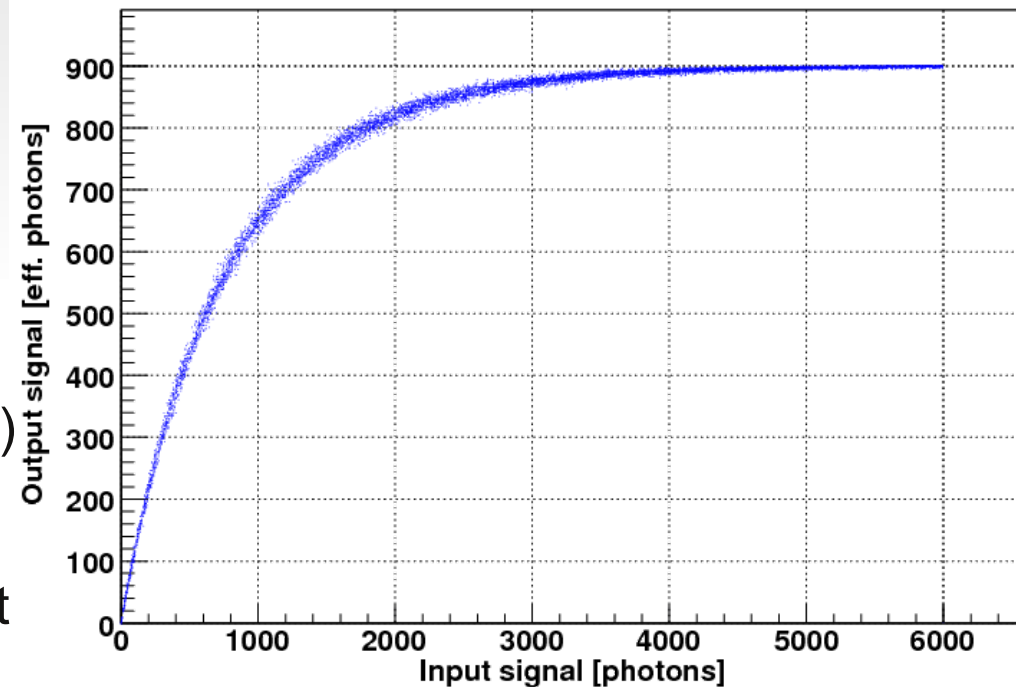


Why G-APDs? saturation



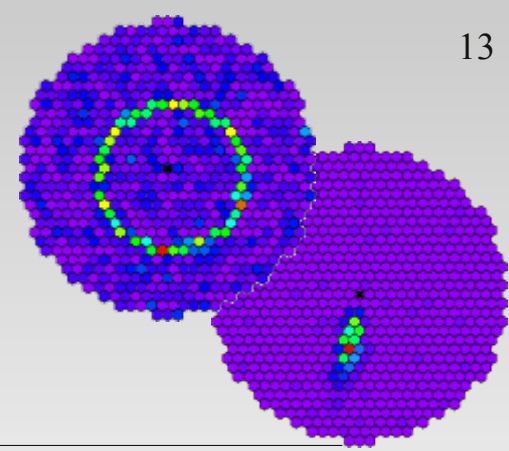
Possible disadvantages:

- **Saturation**
 - ➔ number of cells $\gg R_{\text{nsb}} \cdot T_{\text{recovery}}$
 - ➔ More cells means more dead area (limits the PDE)
- Temperature dependence ($\sim 5\%$ / degree)
 - ➔ temperature **or** gain stabilization via feedback system (stable cal. signal, keep signal height constant by adapting voltage)
- Sensitive $\gg 700\text{nm}$
 - ➔ Increased NSB rate (but NSB rate vs threshold steep enough that energy threshold doesn't change significantly)



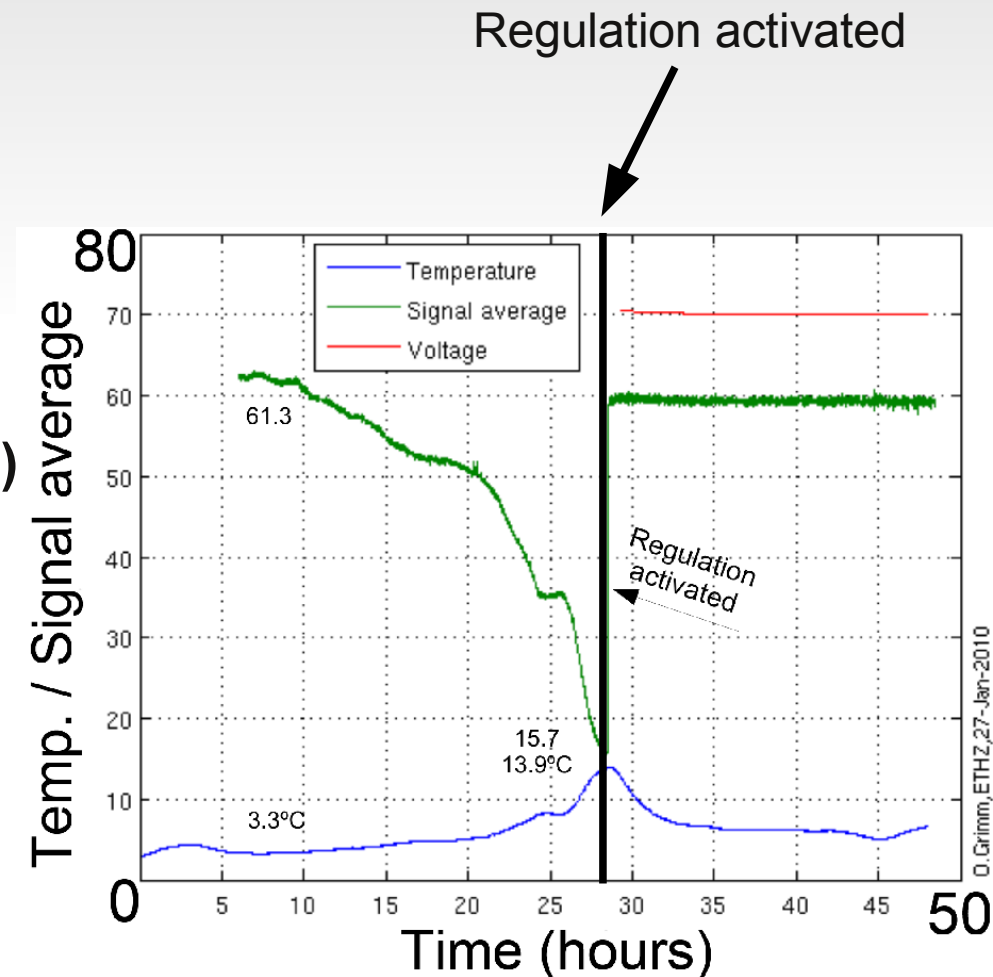


Why G-APDs? temperature dependance



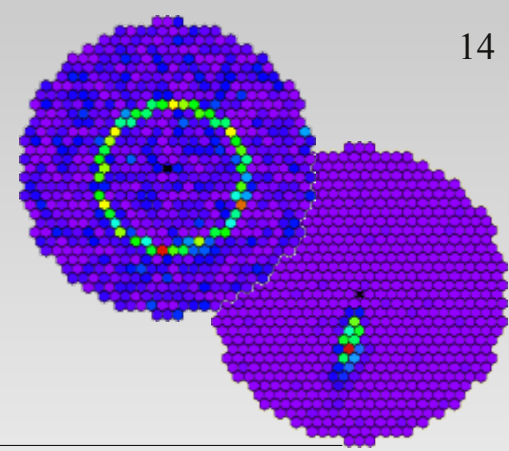
Possible disadvantages:

- Saturation
 - ➔ number of cells $\gg R_{\text{nsb}} \cdot T_{\text{recovery}}$
 - ➔ More cells means more dead area (limits the PDE)
- **Temperature dependance (~5% / degree)**
 - ➔ temperature **or** gain stabilization via feedback system (stable cal. signal, keep signal height constant by adapting voltage)
- Sensitive $\gg 700\text{nm}$
 - ➔ Increased NSB rate (but NSB rate vs threshold steep enough that energy threshold doesn't change significantly)





Why G-APDs?

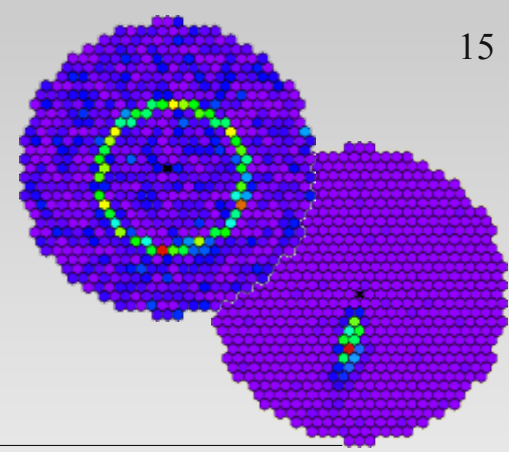


Possible disadvantages:

- Saturation
 - ➔ number of cells $\gg R_{\text{nsb}} \cdot T_{\text{recovery}}$
 - ➔ More cells means more dead area (limits the PDE)
- Temperature dependance ($\sim 5\%$ / degree)
 - ➔ temperature **or** gain stabilization via feedback system (stable cal. signal, keep signal height constant by adapting voltage)
- **Sensitive $\gg 700\text{nm}$**
 - ➔ Increased NSB rate (but NSB rate vs threshold steep enough that energy threshold doesn't change significantly)



Why G-APDs? dark counts

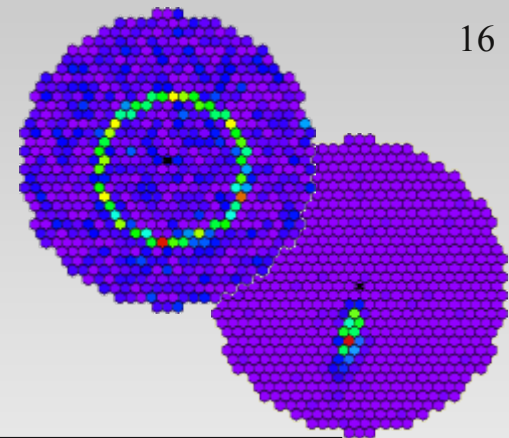


Possible disadvantages:

- Dark counts (in our case $O(5\text{MHz})$ per G-APD)
→ **but NSB rate $O(50\text{MHz})$**
- Afterpulses (gain-dep. prob. 5%-20%)
- Crosstalk (gain-deb. prob. 5%-20%)
- Small collection area
(3 x 3mm, maybe soon 5 x 5mm)



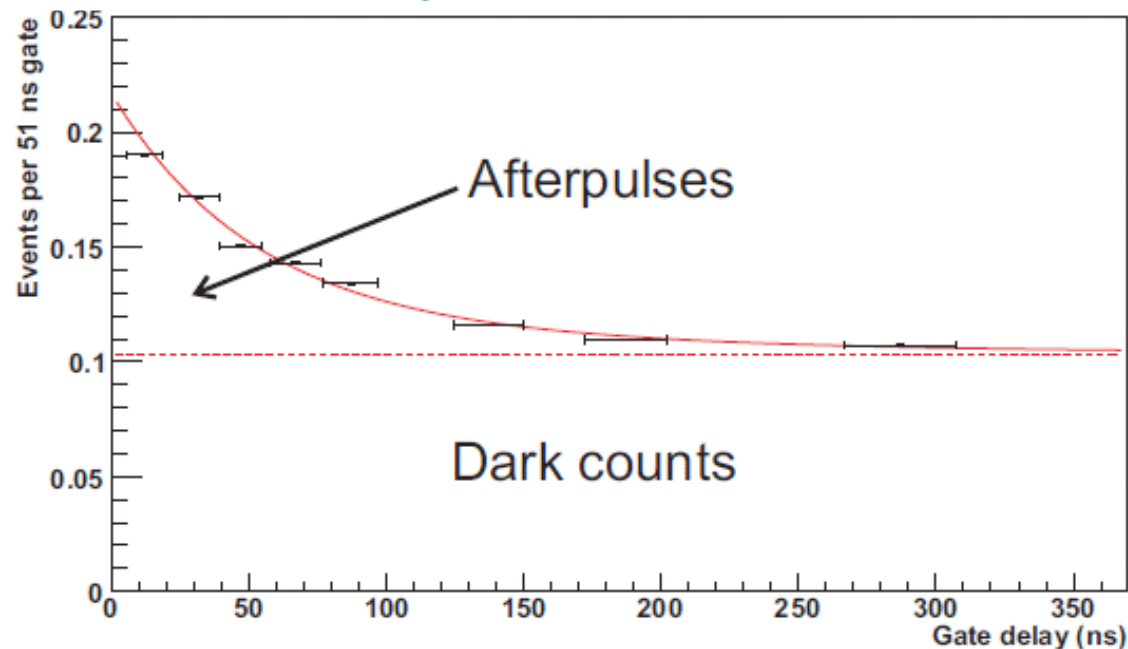
Why G-APDs? afterpulses



Possible disadvantages:

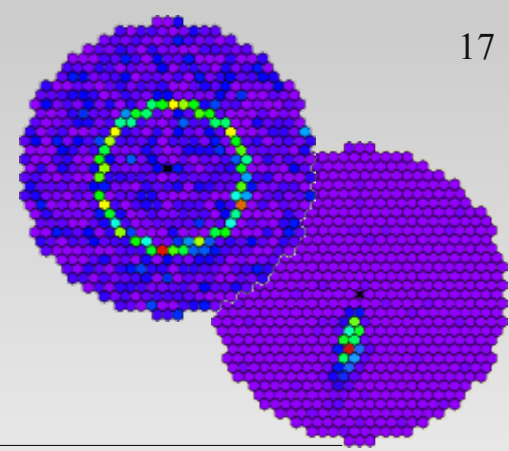
- Dark counts (in our case $O(5\text{MHz})$ per G-APD)
→ **but NSB rate $O(50\text{MHz})$**
- Afterpulses (gain-dep. prob. 5%-20%)
→ **incoherent (prolongate signal, but no fake triggers)**
- Crosstalk (gain-deb. prob. 5%-20%)
- Small collection area
(3 x 3mm, maybe soon 5 x 5mm)

Exponentially decr. prob. after breakdown





Why G-APDs? crosstalk

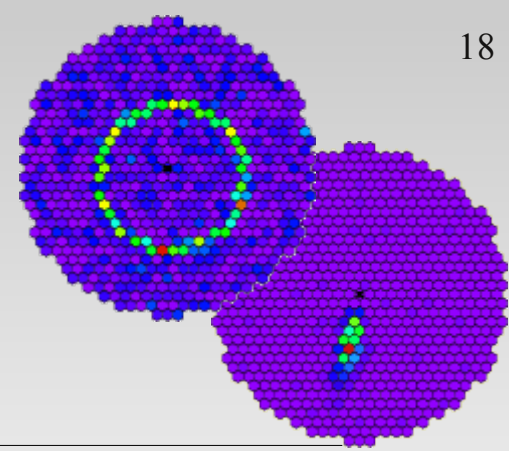


Possible disadvantages:

- Dark counts (in our case $O(5\text{MHz})$ per G-APD)
→ **but NSB rate $O(50\text{MHz})$**
- Afterpulses (gain-dep. prob. 5%-20%)
→ **incoherent (prolongate signal, but no fake triggers)**
- Crosstalk (gain-deb. prob. 5%-20%)
→ **Important for single pe counting, but not for CTs**
→ **just increases the average signal height and slightly its fluctuations**
- Small collection area
(3 x 3mm, maybe soon 5 x 5mm)

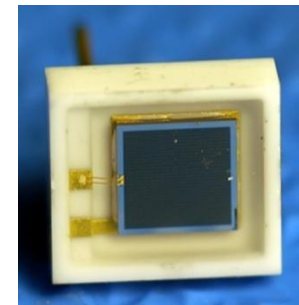


Why G-APDs? collection area



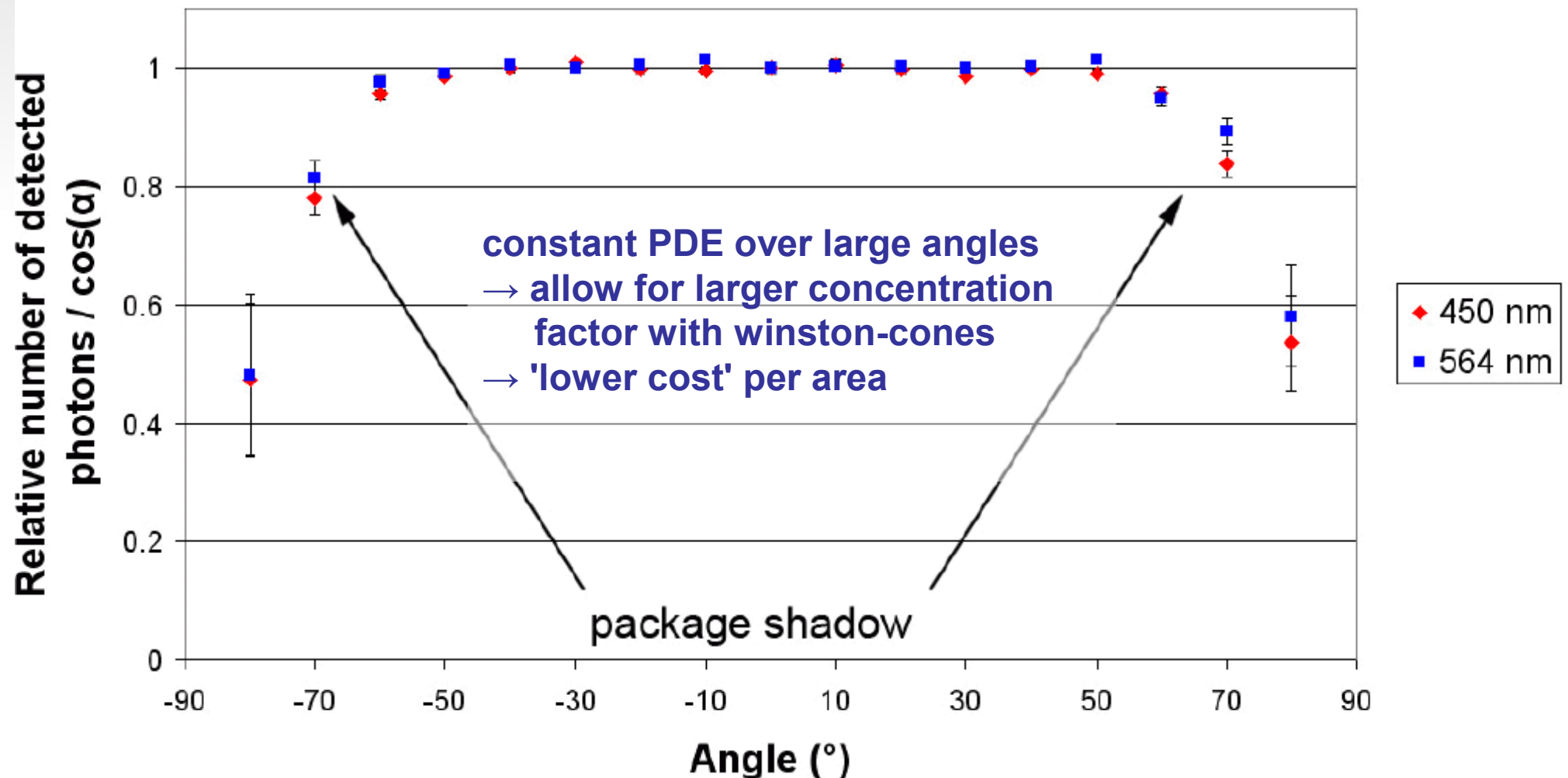
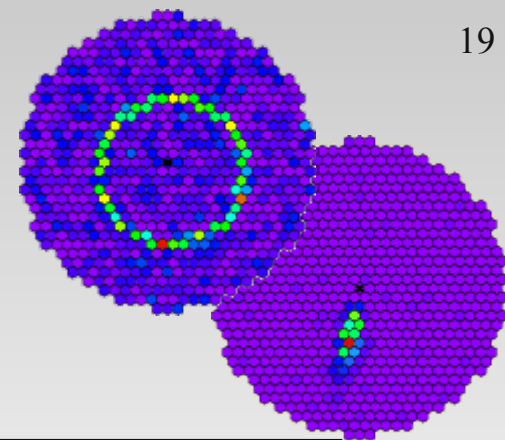
Possible disadvantages:

- Dark counts (in our case $O(5\text{MHz})$ per G-APD)
→ **but NSB rate $O(50\text{MHz})$**
- Afterpulses (gain-dep. prob. 5%-20%)
→ **incoherent (prolongate signal, but no fake triggers)**
- Crosstalk (gain-deb. prob. 5%-20%)
→ **Important for single pe counting, but not for Cts**
→ **just increases the average signal height and slightly its fluctuations**
- Small collection area
(3 x 3mm, maybe soon 5 x 5mm)
→ **But almost $2\cdot\pi$ angular acceptance**



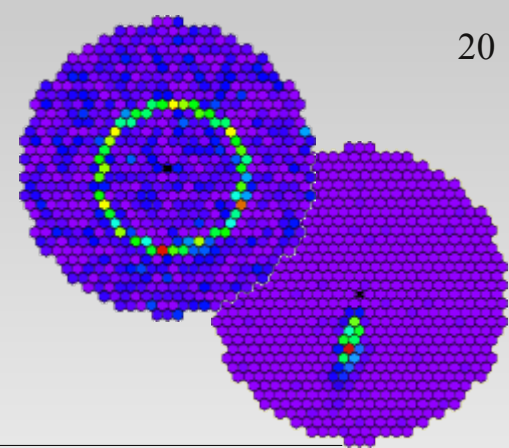


G-APDs angular acceptance



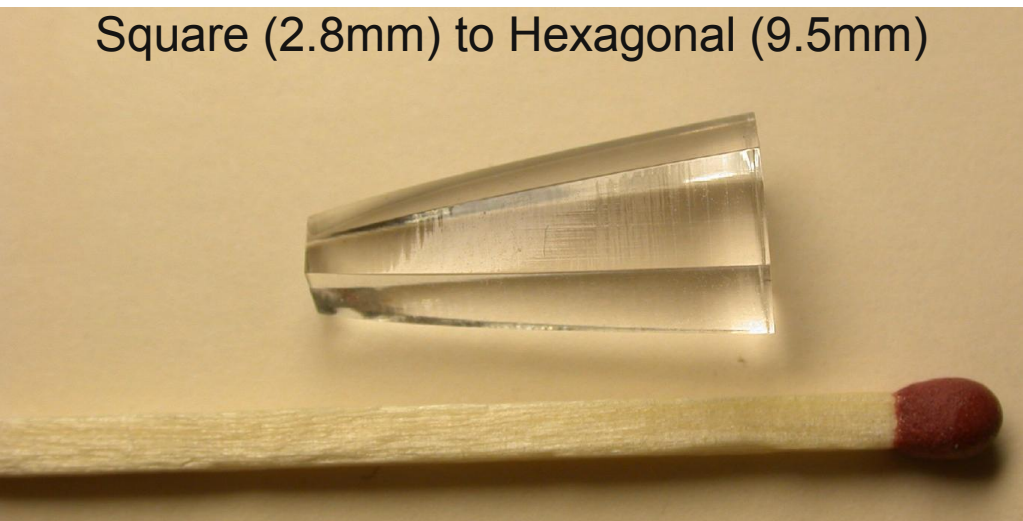


Solid cones - advantages



- The large angular acceptance allows already for a higher compression ratio than for PMTs
- In addition: G-APD properties (small size \rightarrow low weight) allow the use of solid cones

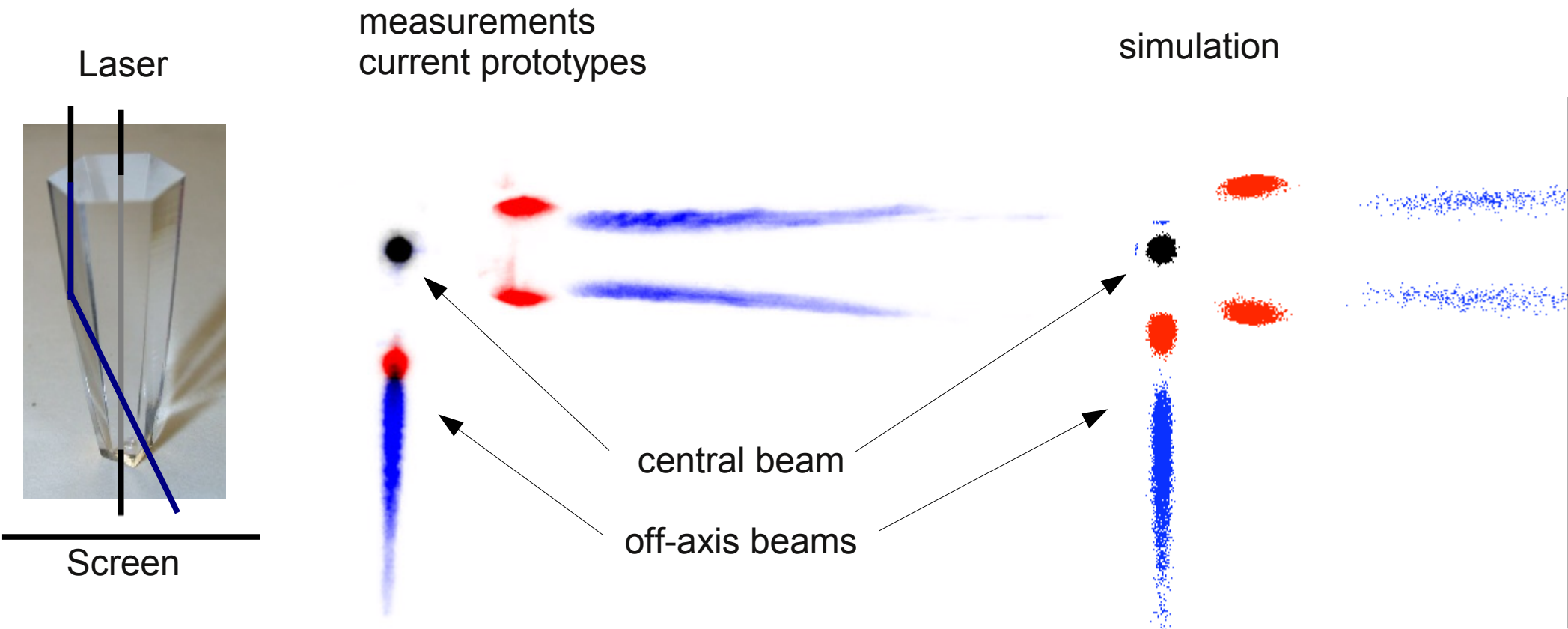
Square (2.8mm) to Hexagonal (9.5mm)



Solid cones

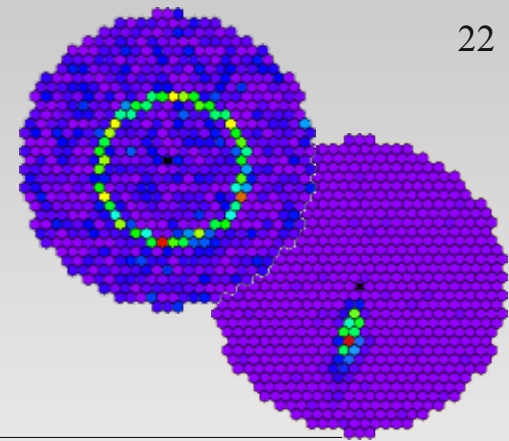
- Higher compression ratio, $O(10)$ (due to Fresnel-reflection)
- Higher reflectivity (total reflection)
- cheaper and easier to produce (injection molding)
- rather complicated shapes possible

Image of laser beam parallel to axis

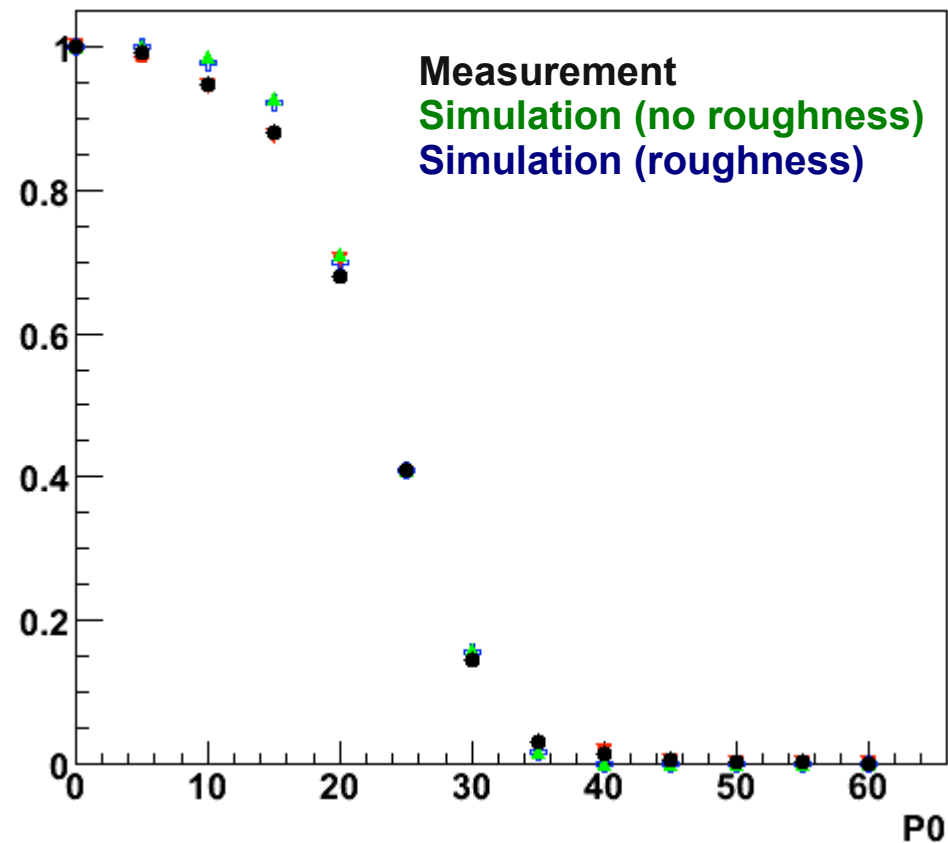




Solid cones angular acceptance

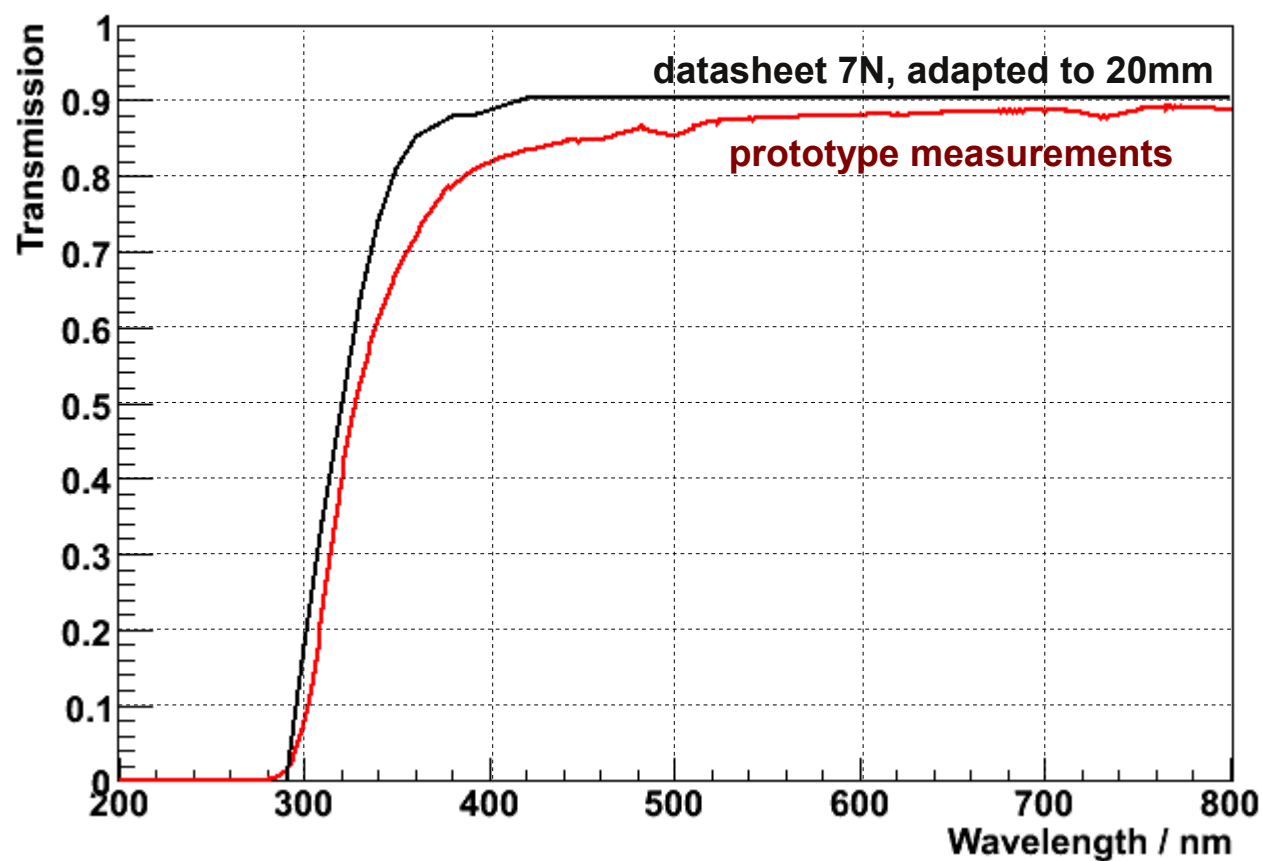
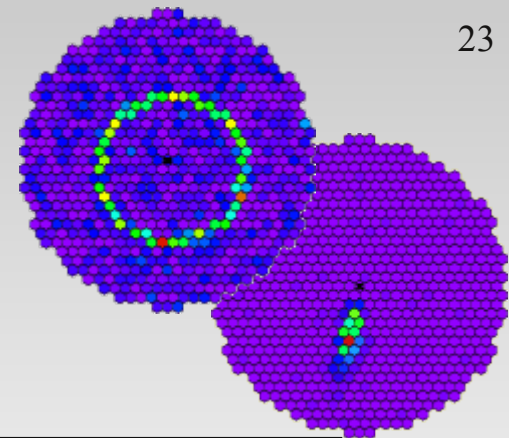


Angular acceptance of measurement
fits the simulation quite well





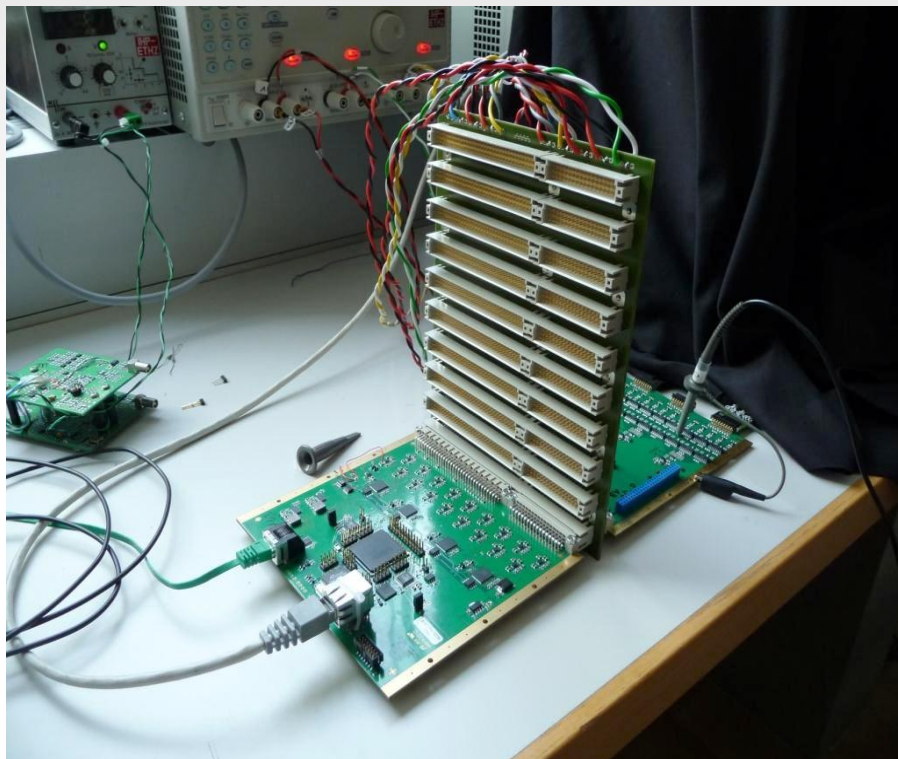
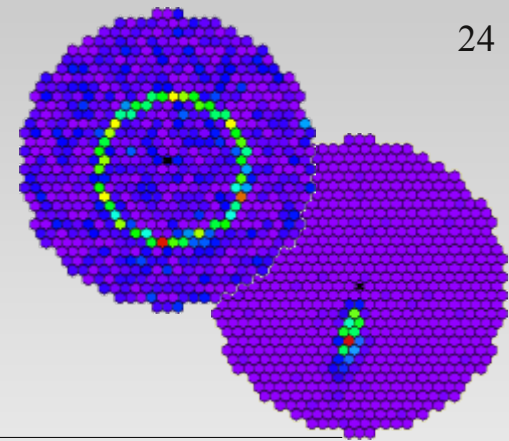
Solid cones - transmission



Light loss compared to data sheet (if folded with C-spectrum and G-APD acceptance) O(%)
 Curves include two Fresnel-reflections (real transmission will be 4% higher)

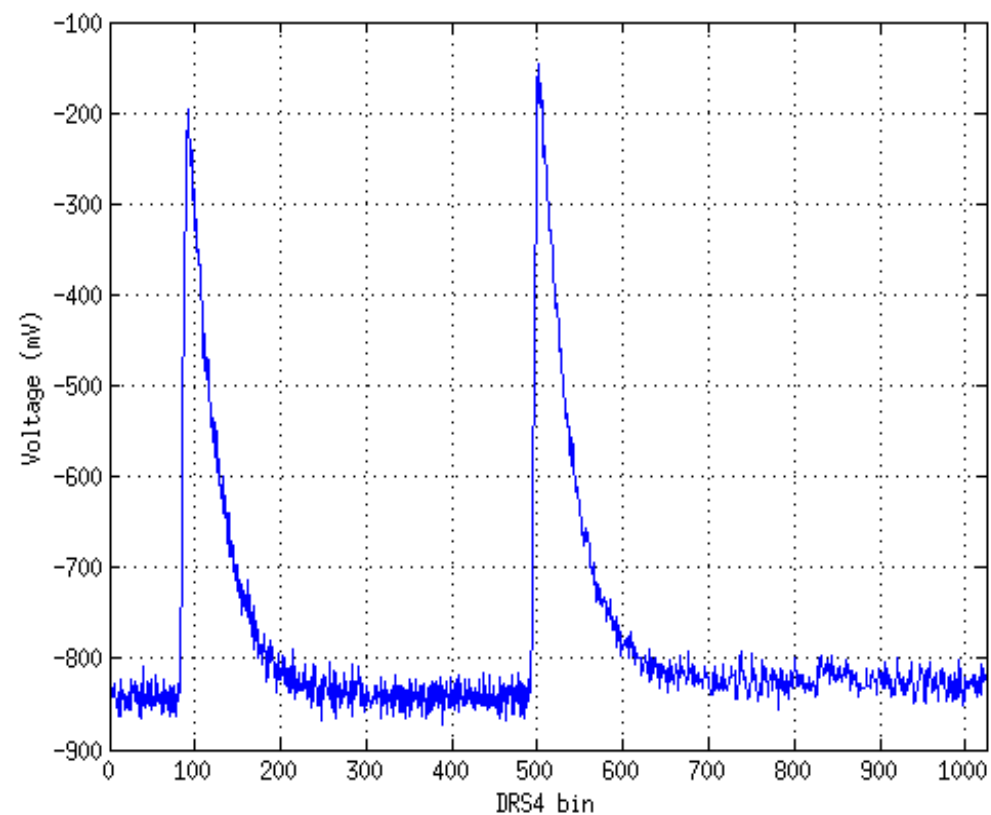


The Camera - prototypes



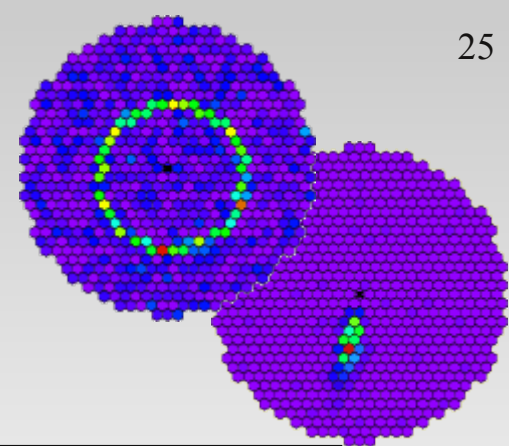
Pre-amplifier board, analog pipeline and digitization board (DRS4) connected via the mid plane, distributing power and slow control signals

LED flashes registered by a G-APD connected to the FACT readout system



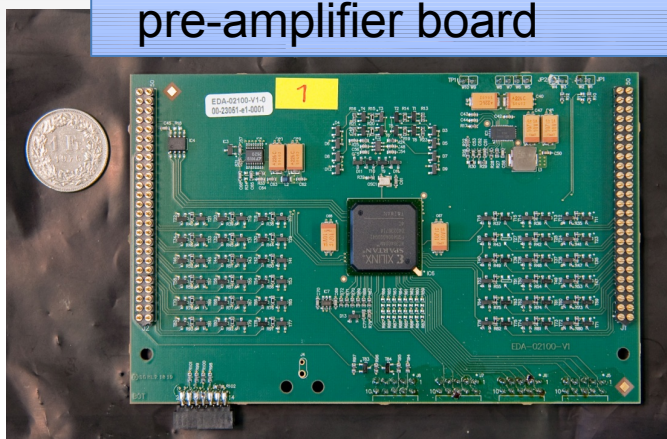


The Camera - prototypes



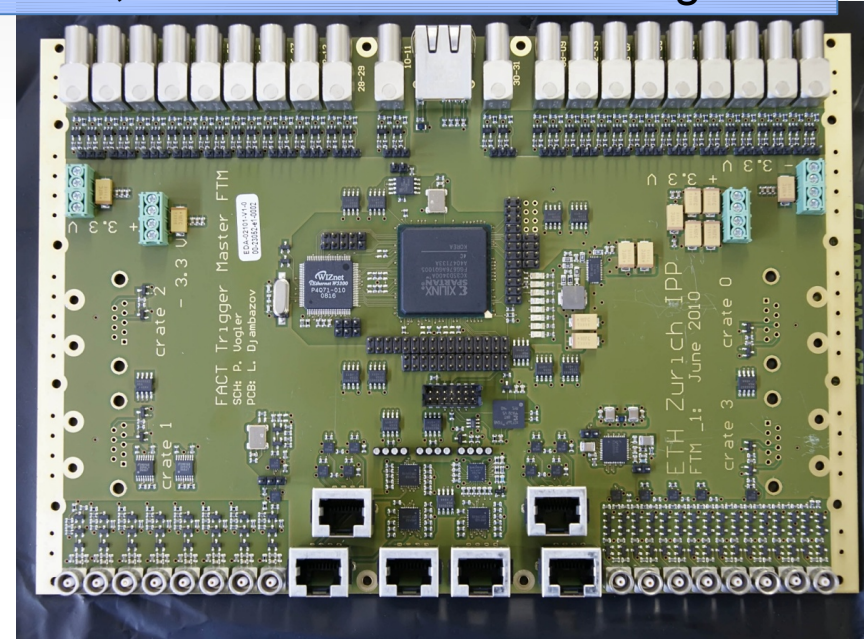
Trigger unit (FTU) – 40 pieces

- Discriminator, rate counting
- Mezzanine card on the pre-amplifier board



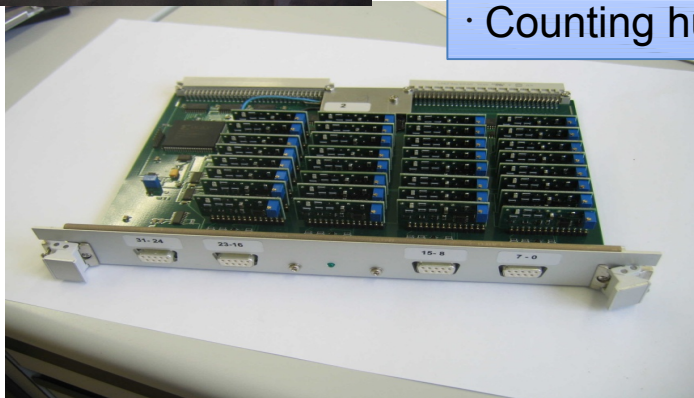
Trigger master (FTM) – 1 piece

- Provides trigger decision upon 40 FTU inputs
- Provides CLOCK, TRIGGER and RESET signals



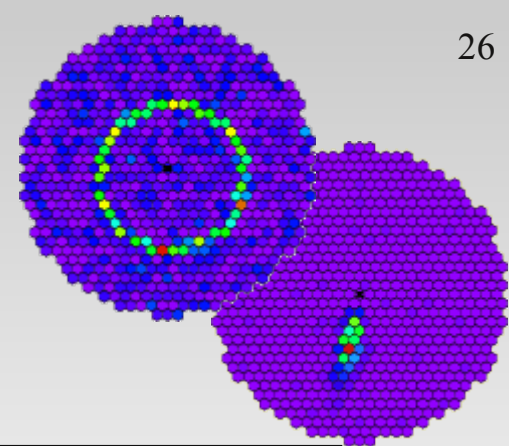
32 channel GAPD bias card

- USB/PC controlled
- Counting hut





Conclusions



- **After two years of hard work, the FACT- telescope will become a reality soon**
 - Mirror and drive system installed
 - Prototypes of major components tested
 - Series production of most components started
 - Readout of a single channel tested (G-APD + pre-amp. + DAQ), gives very promising results
 - Series production of the pre-amplifier boards and readout boards in prep. after minor changes, will start in two weeks
 - Present prototypes for the cones show properties very close to our expectations, production will start soon, too
- ➔ **Full system build ready in the first half of 2011**
- Our experience: **G-APDs seem to be an extremely good alternative to PMTs in IACTs**
 Just a number: You need only ~10 G-APDs with solid cones to get a comparable sensitive area to a MAGIC PMT – 10 x 20€ ~ 200€

Thank you!

Camera: Prototype-electronic available,
starting series production

FACT - First G-APD
Cherenkov Telescope

Mirrors: Installed

Drive: Installed

FACT